

# **Detecting the Geminids Using Radio Waves**

Lucas C.E. Farley

Marlborough College, Marlborough, Wiltshire, SN8 1PA

### Aim

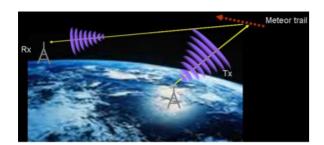
To detect the **Geminids** and their **characteristics** using equipment to receive radio waves at the Marlborough College Blackett Observatory (MCBO). Characteristics include the date and time of the peak and the number of events per hour.

### Introduction

The **GRAVES RADAR** in Dijon, France is a government-run, continuous wave radar which emits radio waves at a 143.050MHz frequency constantly. There are four transmitters at an upward angle of 25°, which each cover a sector of the 180° pointing South and upwards.



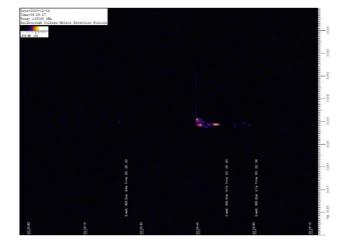
At the MCBO, 690km from GRAVES, we have a 4 element YAGI aerial set up at a 16° angle of elevation which points to around halfway between it and GRAVES. When a meteor or small particle enters the ionosphere around 90-100km above ground, it ablates and ionises the atoms creating positive ions, electrons and neutral molecules in the plasma head of the object and in the ionisation trail. The radio waves intercept both and the electrons reradiate them which is received at the MCBO. The plasma head has a **short** but **strong** signal, the ionisation trail has a long but weak signal, and the overall strength of the signal depends on the **electron density** in both:



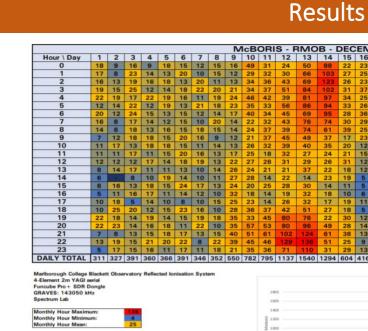
## Method

The 4 element YAGI aerial is connected to a FUNcube **Pro+ Dongle** which is then connected to the **PC** at the observatory. The dongle is tuned to **143.048MHz** and the detection software (Spectrum Lab) is set to show events 1-3kHz above that frequency, so the GRAVES reflections appear around **2kHz above**. This is the series of events which lead to displaying the showers on a graph:

• A significant event is detected and displayed on the computer:



- The x-axis shows time, and the y-axis shows frequency. The volume (in dB) of meteor is shown by its colour with a scale in the top left.
- Blueshift and redshift is shown by the lines going up and down respectively showing frequency increase and decrease. This meteor mostly moving towards us (blueshifted).
- The event must be **loud** enough and **long** enough to be added to the tally of events. This is then added to the heatmap showing the number of events each hour.
- This can then be converted to a bar chart showing events per hour.
- The detection program, shown above, is • constantly running and live. This can be seen at www.blackettobservatory.org



# Conclusion

Detecting the Geminids from the MCBO was successful as our data is very close to the data from NASA:

- The **peak** being between the **4-17**<sup>th</sup> **December**. • Our results show it on the **night of 13**<sup>th</sup>.
- The meteors per hour during the peak is 120.
  - Our results show the highest being **138** at **2200** on 13<sup>th</sup>.

Our data also shows a smaller 2 hour peak on the night of the **12**<sup>th</sup> with 102 events at 2100, and 129 events at 2200. This too approximately matches NASA's data.

### References

NASA. (2019). "Geminids". https://solarsystem.nasa.gov/asteroids-comets-and-meteors/meteors-andmeteorites/geminids/in-depth/

Morgan, D. (2011). "Meteor Radar SDR Receiver (FUNcube Dongle)". https://britastro.org/sites/default/files/MeteorRadarSDRReceiver.pdf Morgan, D. (2011). "Detection of Meteors by RADAR". https://britastro.org/sites/default/files/Detection of meteors by RADAR.pdf





14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31
88	22	23	7	17	9		34	28	29	22	25	16	23	15	20	22	
03	27	25	14	11	15		29	48	21	31	23	24	25	26	19	29	
23	26	23	6	19	11		41	51	28	33	18	18	25	33	30	28	
02	31	37	14	6	14		33	48	46	43	33	30	23	40	32	48	
97	34	25	10	12	18		43	55	32	53	34	31	31	28	33	33	
94	33	26	6	19	15		51	63	43	51	20	30	28	35	24	29	
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22	18	12	14	15	14	31	35	36	33	35	27	23	23	28	33		
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294	604	416	250	276	186	297	803	907	693	677	554	561	543	606	596	279	0

